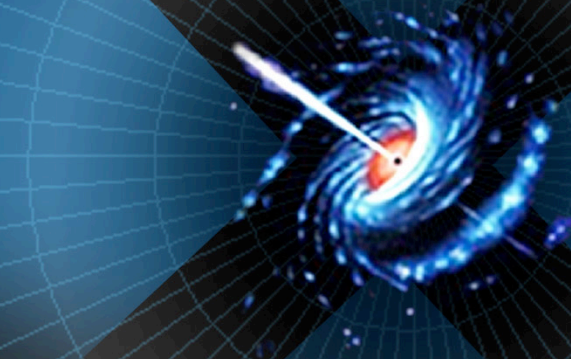


*Constellation-X Facility Science Team Meeting (FST) — February 21/22, 2008*

# Constellation

The Constellation X-ray Mission



## ►► Facility Science Team Panels

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G o d d a r d   S p a c e   F l i g h t   C e n t e r



## **Topics covered : Constellation-X FST Science Panels**

- **Panel restructuring process**
- **Panel topics and panel chairs**
- **Charge to the panels**
- **Logistics/Boundary Conditions**

## **Panel Restructuring Process**

- During Fall 2007 panels re-opened to aid in preparing for the decadal survey (emails sent to high-energy astrophysics community lists)
- More than 120 people responded from >55 institutions including ~13 foreign institutions.
- These 120+ people have been organized into 14 panels: 5 Galactic, 8 extragalactic and one multidisciplinary panel on “Plasma Diagnostics and Atomic Astrophysics”
- An introductory message with a ‘charge to the panels’ was emailed out on 12/20/07
- A telecon with the panel chairs was held January 4, 2008 to discuss the process and prepare for this FST meeting
- Each panel chair has been organizing telecons and/or email exchanges over the last 2 months and will present their work thus far and future plans at this meeting

## Science Panels - Inside the Galaxy

*Note focus is on science themes/topics rather than target type*

| <b>Panel Topic</b>  | <b>Chair</b>                    |
|---|---------------------------------|
| <b>Extreme States of Matter<br/>in Neutron Stars</b>      | <b>Frits Paerels (Columbia)</b> |
| <b>Accretion Physics<br/>in Stellar Systems</b>           | <b>Jon Miller (Michigan)</b>    |
| <b>Production and Distribution<br/>of the Elements</b>    | <b>John Hughes (Rutgers)</b>    |
| <b>MHD Physics<br/>in Stellar Environments</b>            | <b>Rachel Osten (U. of Md)</b>  |
| <b>Solar System, Planet<br/>Formation &amp; Evolution</b> | <b>Eric Feigelson (PSU)</b>     |



## Science Panels - Outside the Galaxy

| <b>Panel Topic</b>  | <b>Chair</b>                           |
|---|--|
| <b>Missing Baryons/WHIM;<br/>synergy with UV spectroscopy</b>                                   | <b>Michael Shull (Colorado)</b>        |
| <b>Census of Black Hole Accretion<br/>in the Universe</b>                                       | <b>Nancy Levenson (U. of Kentucky)</b> |
| <b>Evolution of Large Scale Structure<br/>in the Universe</b>                                   | <b>Steve Allen (KIPAC/Stanford)</b>    |
| <b>Hot Baryons in Deep Potential Wells</b>  | <b>Christine Jones (SAO)</b>           |
| <b>Testing General Relativity and<br/>Measuring Black Hole Spin</b>                             | <b>Chris Reynolds (U. of Md)</b>       |
| <b>Supernova/Stellar Feedback</b>   | <b>David Strickland (JHU)</b>          |
| <b>AGN Feedback: Outflows &amp; Jets</b>  | <b>Andy Fabian (Cambridge)</b>         |
| <b>The high-z Universe, Re-ionization &amp;<br/>Synergy with JWST "First Light"<br/>Science</b> | <b>Niel Brandt (PSU)</b>               |

## Science Panels - Cross-Panel Working Group on Diagnostics

| Panel Topic                                       | Chair                         |
|---|-------------------------------|
| <b>Plasma Diagnostics and Atomic Astrophysics</b> | <b>Nancy Brickhouse (SAO)</b> |

## Charge to the Panels

*The Science Panels are expected to review, update, and strengthen the Constellation-X science case in preparation for the upcoming decadal survey.*

## Why now? (beyond the decadal survey timing)

- Focus the science case back to being an “astrophysics observatory” rather than a Beyond Einstein mission
- The last science case overhaul was in late 2004/early 2005: the mission configuration has changed since then
  - XGS has higher spectral resolving power ( $R=1250$  is the requirement)
  - XMS has a larger FOV ( $5' \times 5'$  versus  $2.5' \times 2.5'$ )
  - HXT has reduced collecting area ( $150 \text{ cm}^2$  at 40 keV)

## Work Flow for the Science Panels

1. Highlight one or more key science objectives within the topical area
  - Generate some text motivating why this is important
    - broad astrophysics perspective
    - why X-rays are unique or essential to the science
2. Identify possible measurements/observations with Con-X which can be evaluated through simulations.
3. Produce science materials for promoting Constellation-X:
  1. Brief, 'elevator-pitch' summaries including Power Point slides
  2. Detailed simulations demonstrating how the science goals map to the mission parameters.
  3. Written text (TBD): ApJ papers and/or conference proceedings are encouraged

## Work Flow for the Science Panels

**February  
2008**

1. Highlight one or more key science objectives within the topical area

- Generate some text motivating why this is important
  - broad astrophysics perspective
  - why X-rays are unique or essential to the science

**Spring  
2008**

2. Identify possible measurements/observations with Con-X which can be evaluated through simulations.

3. Produce science materials for promoting Constellation-X:

**Finish by  
August  
2008**

1. Brief, 'elevator-pitch' summaries including Power Point slides
2. Detailed simulations demonstrating how the science goals map to the mission parameters.

3. Written text (TBD): ApJ papers and/or conference proceedings are encouraged

**Fall  
2008**



## Logistics/Boundary Conditions

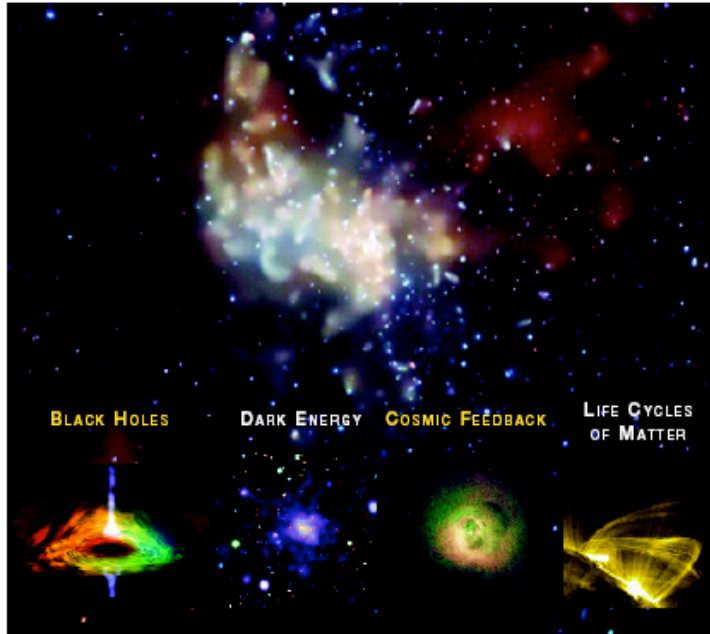
- There is a new email address that reaches all three of Randall, Mike and Ann: [conx-observer@lists.nasa.gov](mailto:conx-observer@lists.nasa.gov)
- **Mission parameters:**
  - Mission cost and programmatic risk are expected to be factors in the decadal survey process.
  - ANSWER THIS QUESTION:  
What science can be done with the mission as described?  
(for now, imagine you are writing proposals, not designing the mission)
  - For science simulations use the parameters as defined for the Atlas V launch vehicle configuration : the latest response matrices are the FEB 2008 versions at [constellation.gsfc.nasa.gov](http://constellation.gsfc.nasa.gov)
  - Also please use the new simx simulator (R. Smith's talk, NEXT)

**Thank you!**

NASA/TP-2005-212784



### Science with Constellation-X



May 2005

- We hope to be producing great new products this fall based on your hard work.
- Thank you VERY much for taking the time to participate in the Con-X FST Science Panels

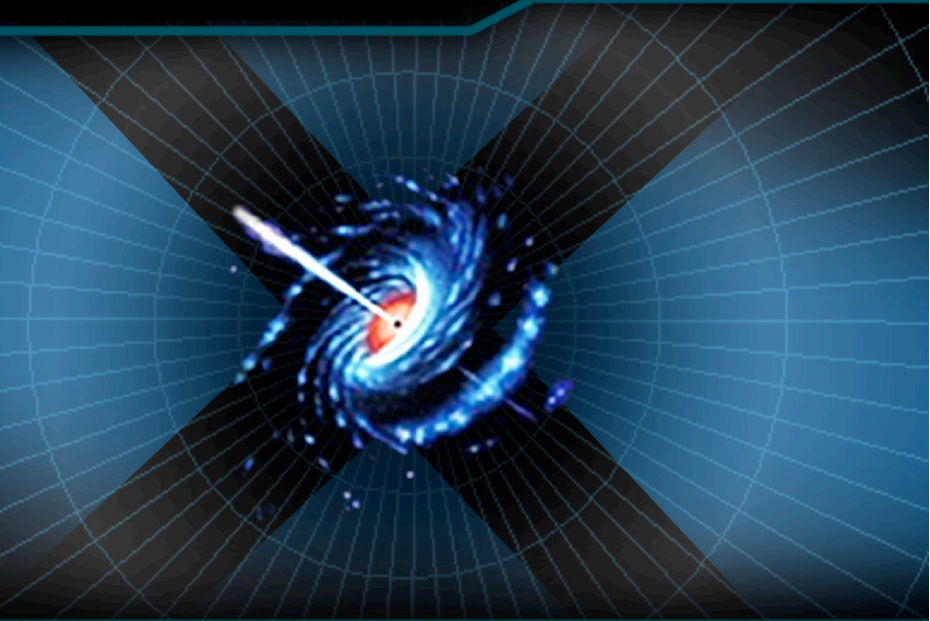
## Key Performance Requirements

|                      |   |
|----------------------|---|
| Effective Area:      | 15,000 cm <sup>2</sup> @1.25 keV<br>6,000 cm <sup>2</sup> @6 keV<br>150 cm <sup>2</sup> @40 keV |
| Bandpass:            | 0.3 – 40 keV  |
| Spectral Resolution: | 1250 @0.3 – 1 keV<br>2400 @6 keV  |
| Angular Resolution   | 15 arcsec 0.3 – 7 keV<br>30 arcsec 7.0 – 40 keV   |
| Field of View        | 5 x 5 arcmin  |

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# Constellation

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▶▶ **simx:**

**The Constellation-X Event Simulator**

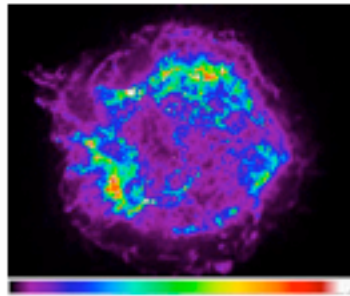
*Randall Smith*

*NASA/GSFC & JHU*

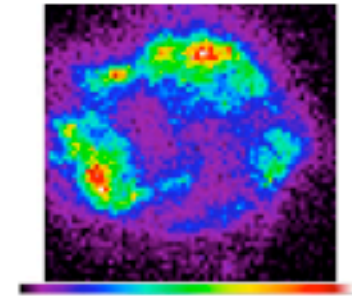
G o d d a r d   S p a c e   F l i g h t   C e n t e r



<http://lheawww.gsfc.nasa.gov/users/rsmith/simx/index.html>



**simx: A  
Constellation-X  
Event  
Simulator**



[Download simx](#)

[Install simx](#)

[Run simx](#)

#### Downloading simx

Click to download the tarball for [simx](#). The current version is 0.2. The model psf is based on the simulated performance on-axis at 6.0 keV; no variation with energy or off-axis position is included. Energy resolution is based on the current response matrix. **Simx** is not particularly fast; we are working

#### Installing simx

**simx models the SXT PSF and XMS response, generating OGIP-standard event files**

## Installation

- Download simx from

<http://hea-www.gsfc.nasa.gov/users/rsmith/simx/>

- Install via: (requires c compiler)

```
unix% tar -zxf simx-0.2.tar.gz
```

```
unix% cd simx-0.2
```

```
unix% ./configure
```

```
unix% make
```

```
unix% make install
```

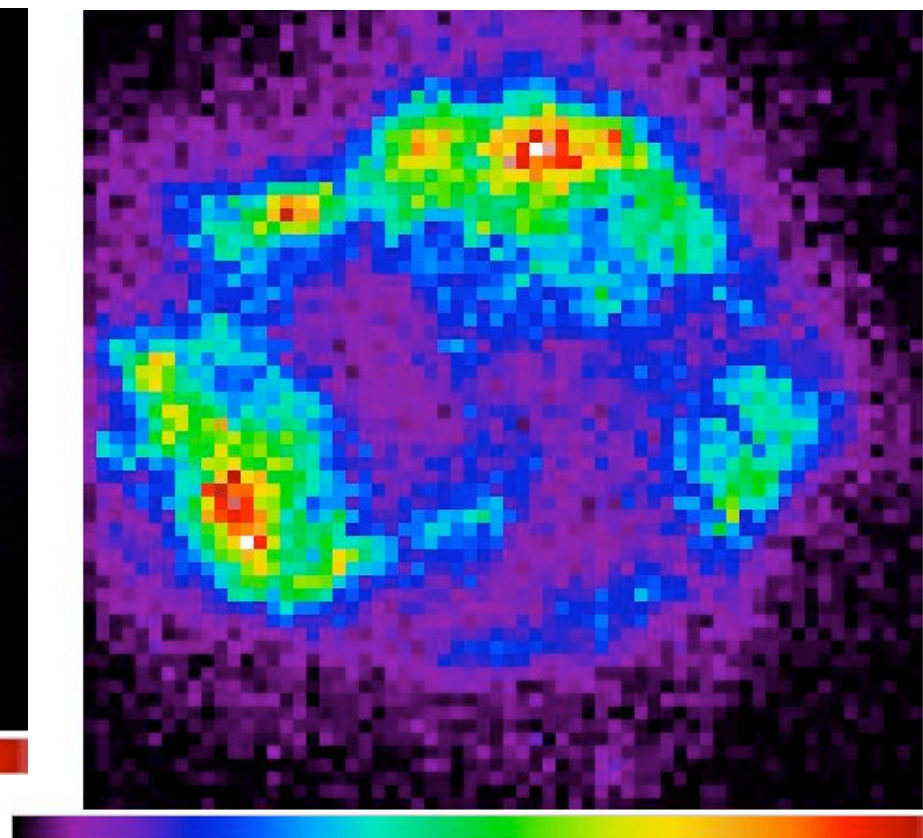
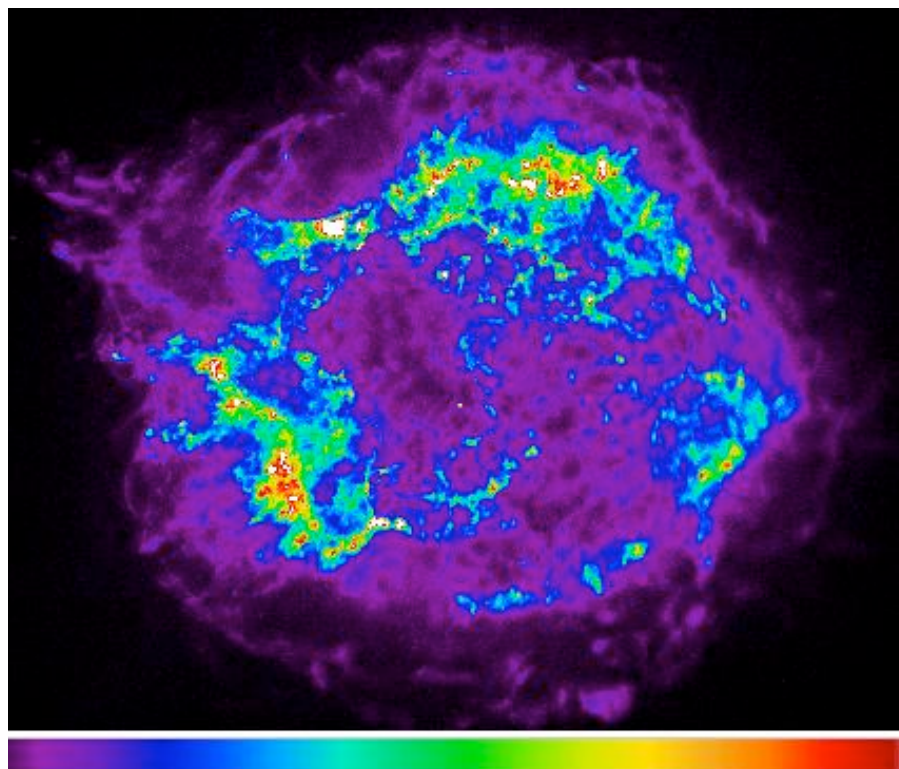
- Tested on OS X, various flavors of Linux, and Solaris



## Running simx

- Detailed instructions in the 'doc/' directory list how to install and run simx.
- Sample run:
  - `unix% pset simx OutputFileName=MyRun`
  - `unix% pset simx Exposure=100000`
  - `unix% pset simx SourceFlux=5e-11`
  - `unix% pset simx SourceImageType=Image`
  - `unix% pset simx SourceImageFile=a2029_img.fits`
  - `unix% pset simx SourceSpectrumType=spec.dat`
  - `unix% pset simx SourceSpectrumTpe=XSPEC_Fiel`
  - `unix% simx`
- Output is MyRun\_evt.fits, usable with ds9, xselect (with some assistance) or dmextract

## Sample Output



Cas A seen with Chandra and Con-X

## **simx Features and Plans**

- **Uses either point source or image.**
  - Goal to include a ‘photon-list’ input including both position and energy for each photon.
- **SXT PSF based on P. Reid simulation on-axis at 6 keV, runs to 100” at maximum.**
  - Goal to include position & energy-dependent PSF
- **For bright sources, simx estimates deadtime at runtime.**
  - Goal to flag ‘degraded’ events in event files.

## Possible Features

- Include vignetting model
- Include background terms from:
  - Instrumental/Cosmic Ray effects (Non-X-ray Bgd)
  - Cosmic X-ray Background
  - Solar-wind Charge Exchange effects
- Modeling XMS pixel differences; central 2.5' pixels may have higher resolution than outer pixels.
- Web-based version of simx (note that input and output files can be large)
- Other suggestions?

## Conclusions

- Available (currently) at <http://lheawww.gsfc.nasa.gov/users/rsmith/simx>
- Will soon move to the main Con-X website.
- Although basic, simx is adequate to determine how well a point source can be extracted from diffuse emission or other sources, or to see how well a radial profile can be measured.
- Contact [Randall.K.Smith@nasa.gov](mailto:Randall.K.Smith@nasa.gov) with suggestions for improvements or bug reports...